ASTR 340: Origin of the Universe

Prof. Benedikt Diemer

Lecture 11 • The Universe beyond our Galaxy

10/05/2021

Homework #1

| ASTR340 > Assignments | |
|-----------------------|--|
| Fall 2021 Home | Homework #1 Due: Thu Sep 23, 2021 11:59pm |
| Syllabus | Attempt 1 V Next Up: Submit Assignment |
| People | |
| Assignments | Unlimited Attempts Allowed |
| Discussions | Available: Sep 9, 2021 1:45pm until Sep 23, 2021 11:59pm |
| Quizzes | ✓ Details |
| Clickers | |
| Grades | Please see the homework 1 pdf file for the questions. For your solution, please submit a pdf file, which you can scan from hand-written pages or create digitally. |
| Zoom | You can find the solutions here. |
| Panopto Recordings | |

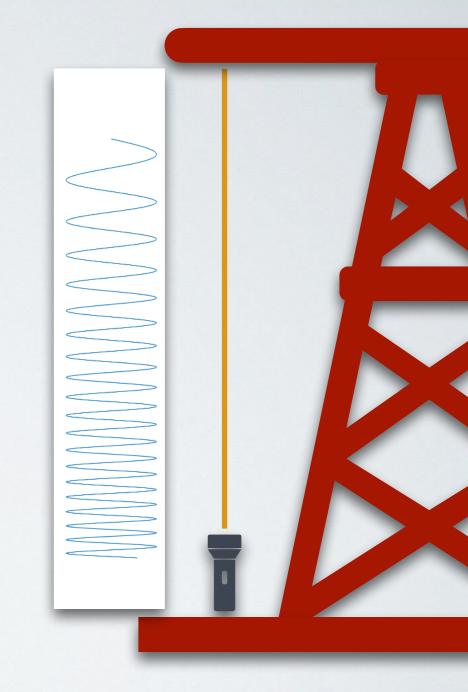
Quiz grades

- No adjusting quiz grades after the fact... sorry!
- Average scores are generally well above 90%
- Two attempts
- Worst grades are not dropped (as for attendance)

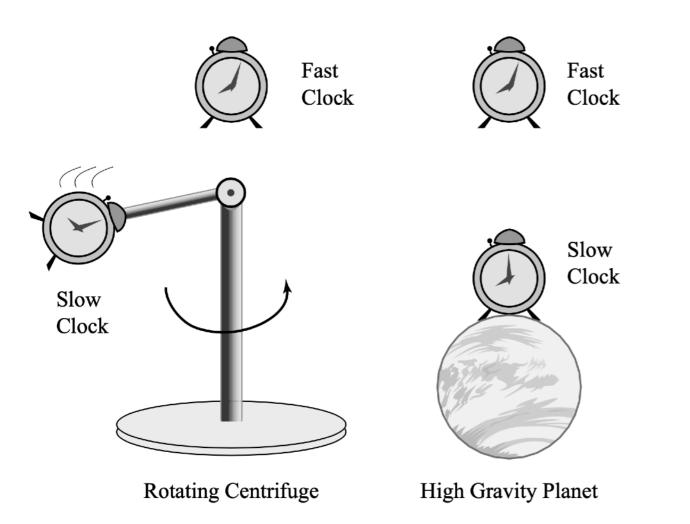
Gravitational time dilation

- Light beam loses energy as it climbs up (gravitational redshifting)
- Frequency decreases
- Imagine a clock based on frequency of laser light: 1 tick = time taken for fixed number of crests to pass
- Gravitational redshifting slows down the clock
- Clocks in gravitational fields run slower

$$\Delta t_{\rm grav} = \sqrt{1 - \frac{2GM}{c^2 r}} \Delta t_{\rm space}$$



Alternative reasoning: spinning clock



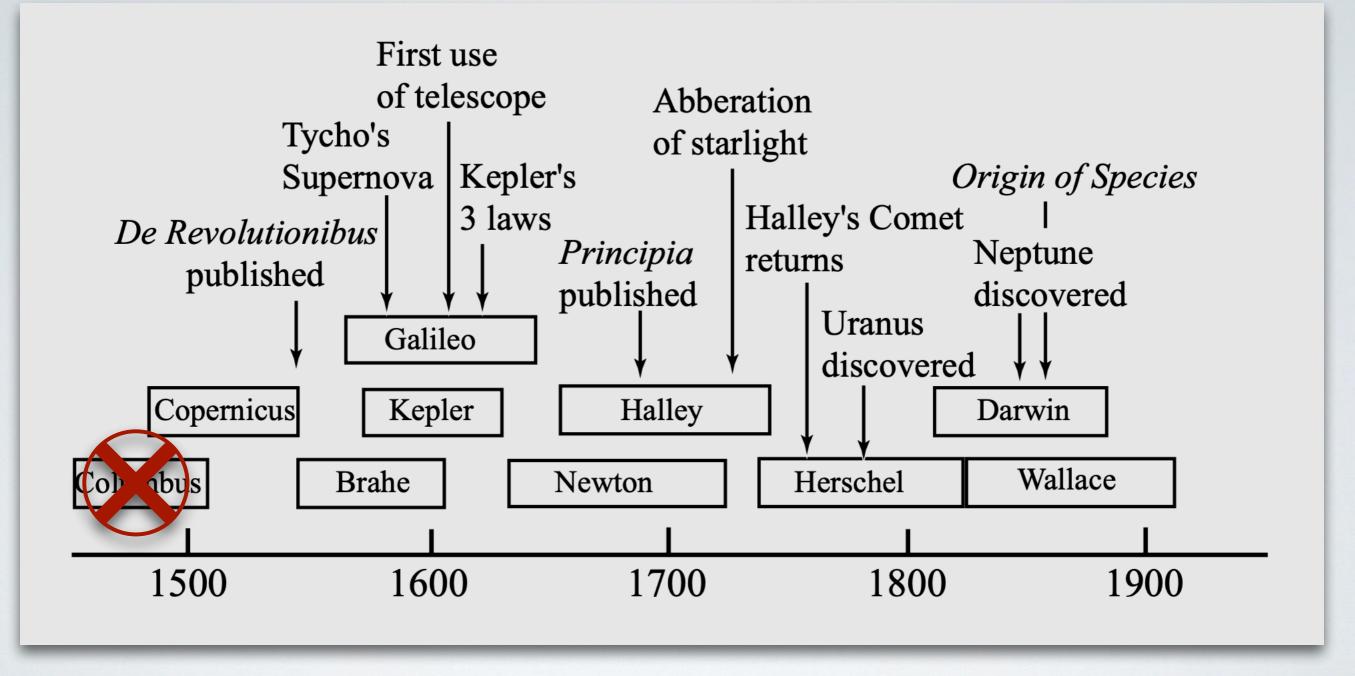
- Imaging quickly rotating clock
- Not an inertial frame, so SR doesn't apply directly
- But as it turns out, a stationery observer still sees the clock running more slowly than in its own (proper) frame
- The circular rotation is equal to a constant acceleration
- According to the weak equivalence principle, this acceleration is indistinguishable from gravity
- Thus, clocks must run more slowly in gravitational fields

Today

- Enlightenment Science
- The Great Debate
- The distance ladder
- Hubble & the expanding Universe

Part 1: Enlightenment Science

Today



Hawley & Holcomb, Foundations of Modern Cosmology



TurningPoint: At the beginning of the enlightenment era, do we know the scale of the solar system?





TurningPoint: At the beginning of the enlightenment era, do we know the distance to the closest stars?





TurningPoint: At the beginning of the enlightenment era, what do we know about the Milky Way?

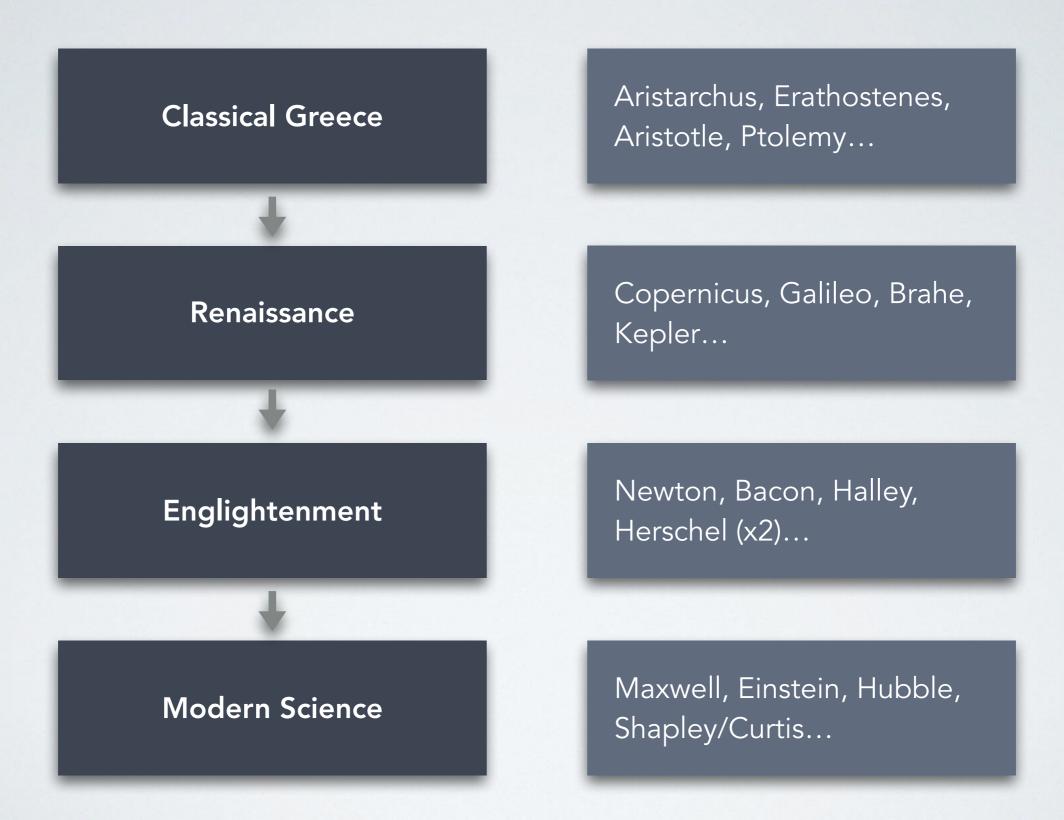




TurningPoint: At the beginning of the enlightenment era, what do we know about the Universe as a whole?



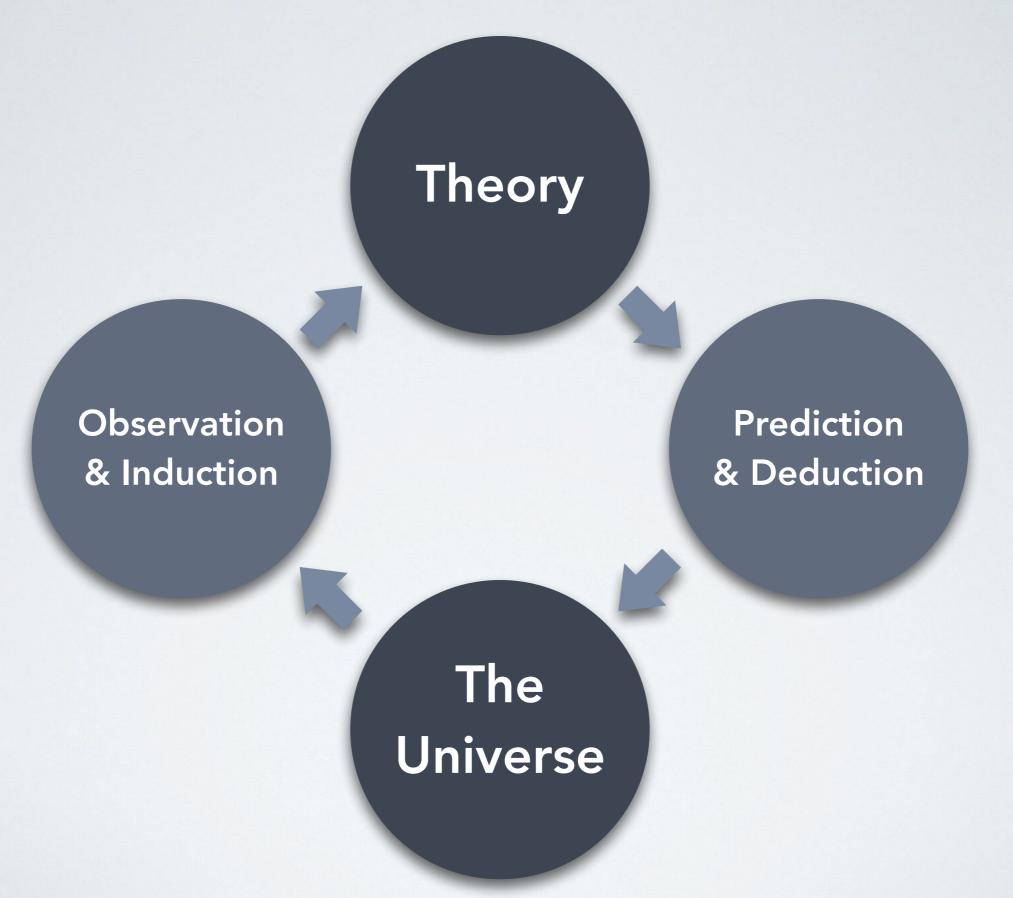
History



Enlightenment science

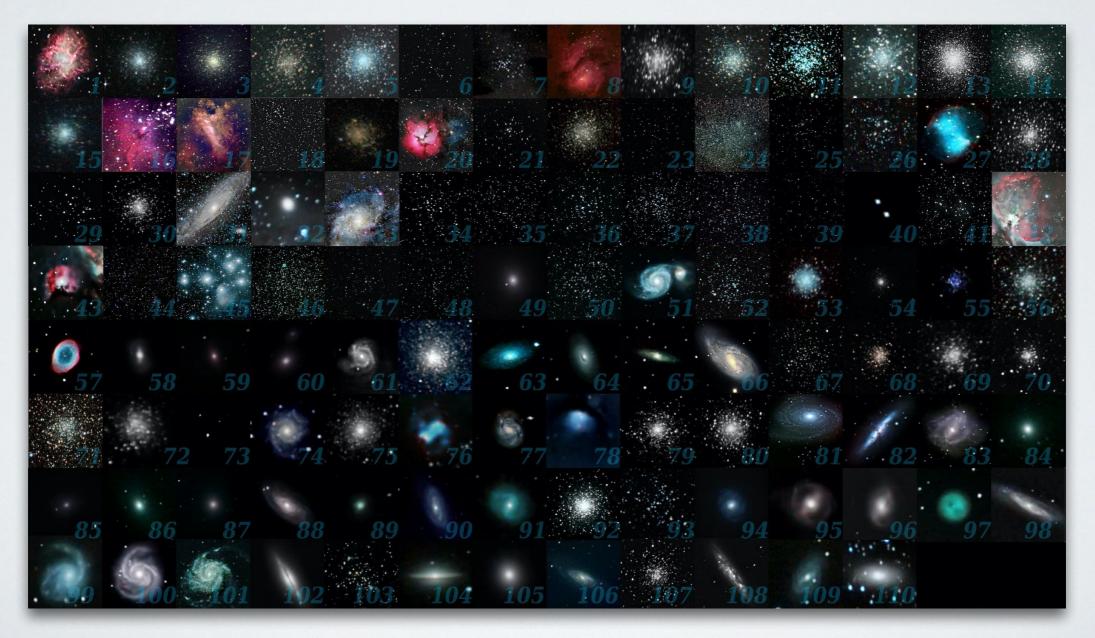
- Beginning of enlightenment often credited to...
 - Bacon's scientific method (~1620)
 - Newton's Principia (~1687)
- Important enlightenment philosophers (for this course):
 - Francis Bacon
 - Immanuel Kant
- Important enlightenment scientists (for this course):
 - Messier
 - Halley
 - The Herschels

The scientific method

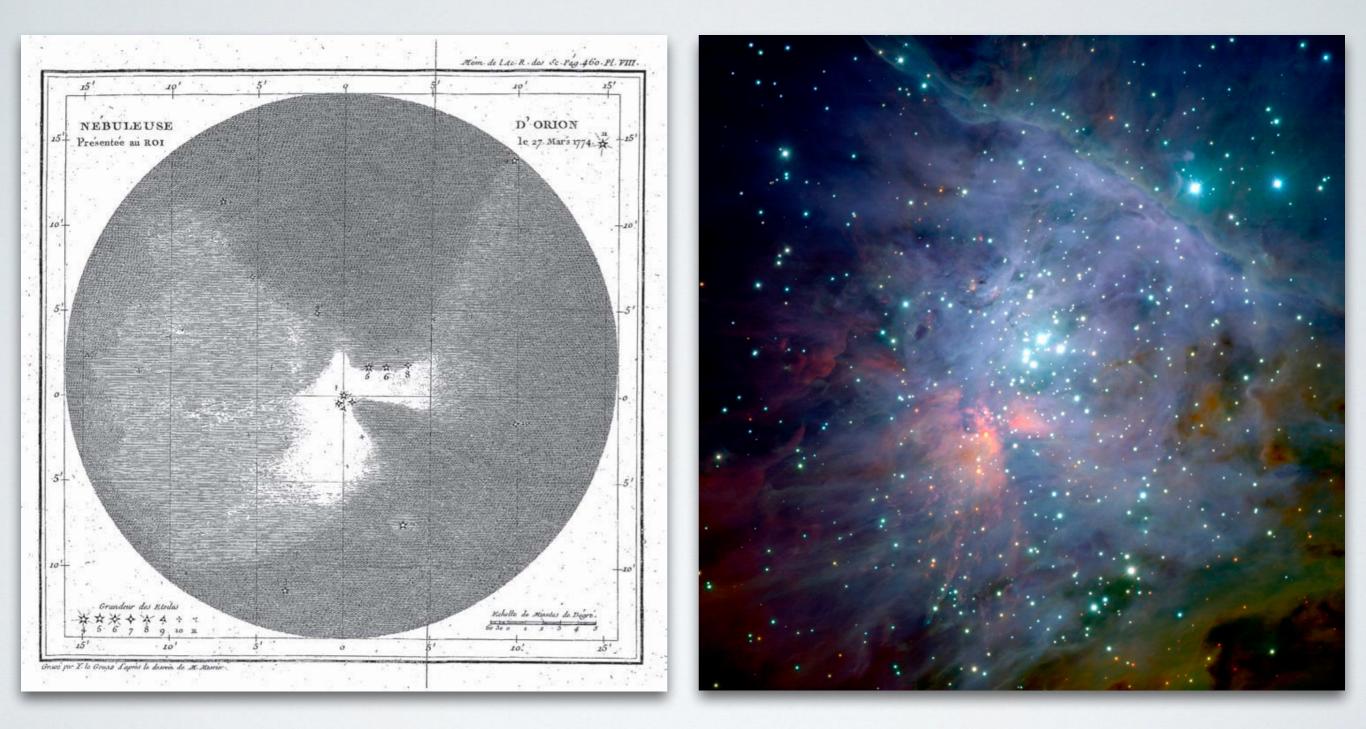


Charles Messier (1730 - 1817)

- Identified many **nebulae** (fuzzy patches of light)
- Published Messier Catalogue in 1780
- Intended as aids to **comet hunters** to reject "uninteresting" objects
- Catalog contains galaxies, star-forming clouds, star clusters...

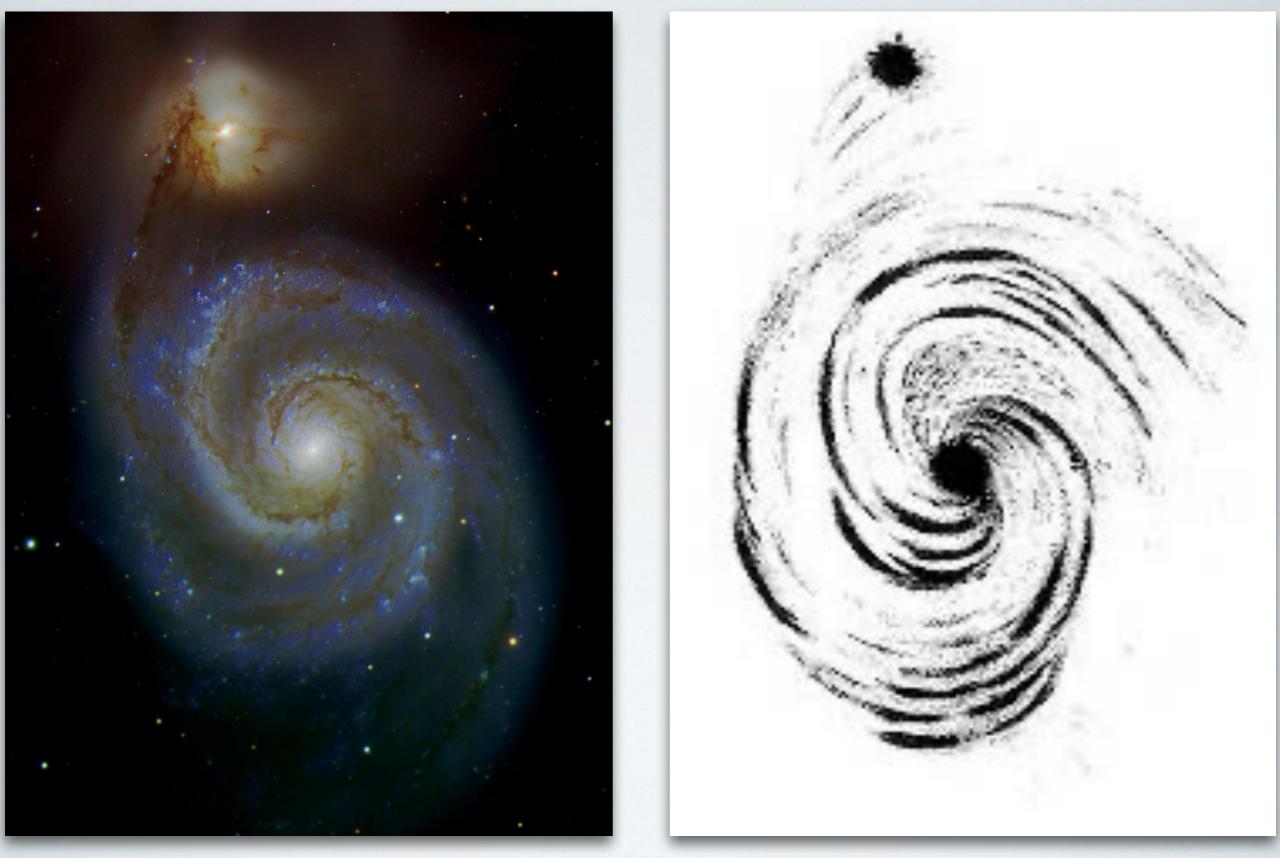


Orion Nebula (M42)



Images: ESO / Wikipedia

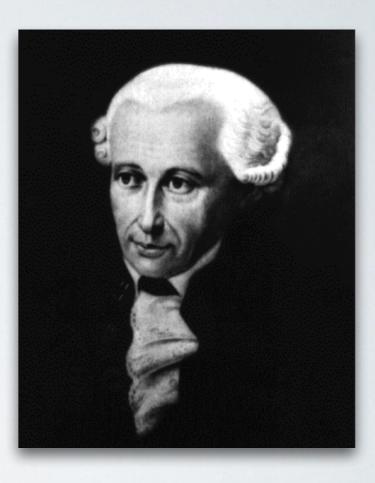
Whirlpool Galaxy (M51 a/b)



Drawing by Lord Rosse (1845)

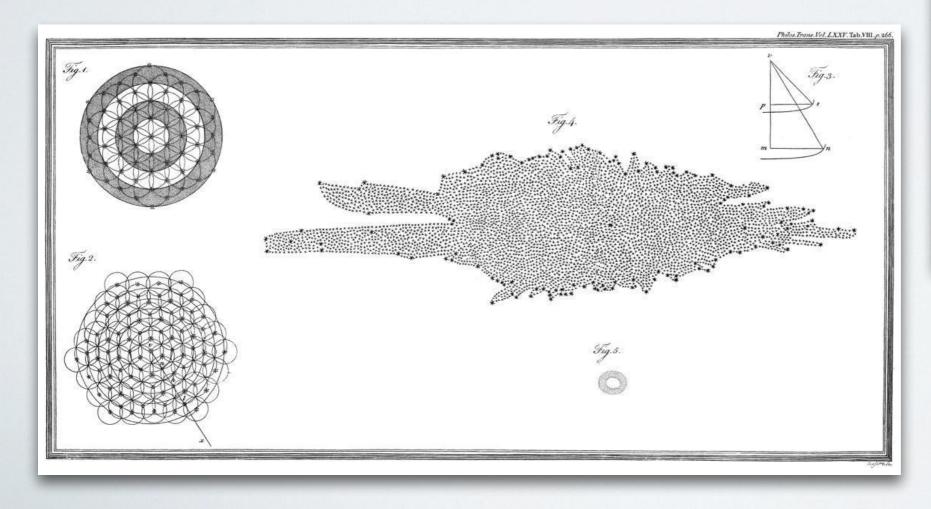
Immanuel Kant and the Island Universe

- Philosopher but also interested in astronomy
- Heard about the "nebulae" and postulated that they are separate "worlds" similar to Milky Way
- Called them "Island Universes" (1775)
- Views did not take hold becaues there was no direct evidence yet of other galaxies



William & Caroline Herschel

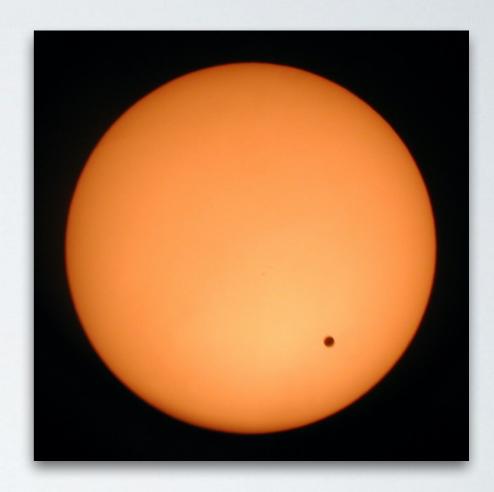
- Discovered Uranus (1781)
- Discovered about 2400 new nebulae
- Thought they could be star collections like MW
- Concluded Milky Way was a disk with the Sun at the center (map from 1785)





Size of the Solar System

- Many attempts to improve on Aristarchus' measurement
- Gregory / Halley proposed using (rare) transit of Venus (as seen from different points on Earth)
- Measured two transits in 1761/1769
- Early example of **collaboration** of many astronomers!
- Got solar parallax of 8.6" (within 4%)

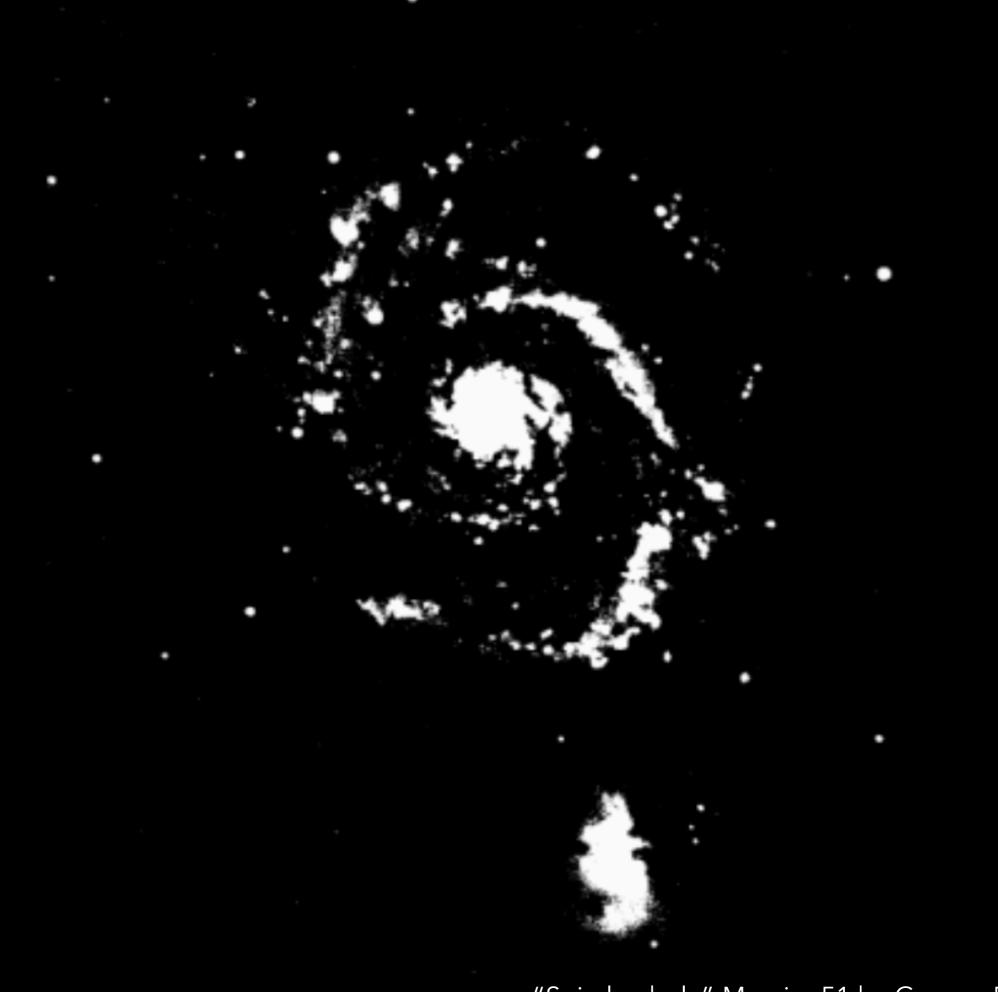


Enlightenment Summary

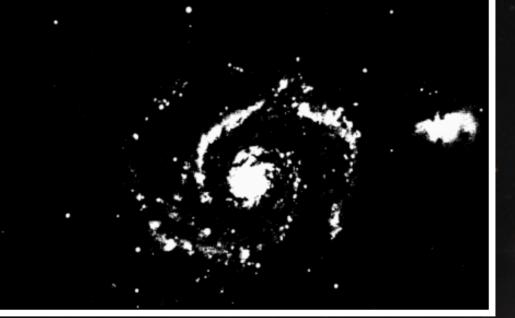
• After the enlightenment (beginning of 19th century), we...

- have observed thousands of nebulae with different shapes
- know the scale of the solar system
- roughly know the shape of Milky Way
- but...
 - do not know whether nebulae are in MW or "island universes"
 - do not have a good scale for the Galaxy (no stellar parallax)
 - think the Universe is static

Part 2: The Great Debate of 1920

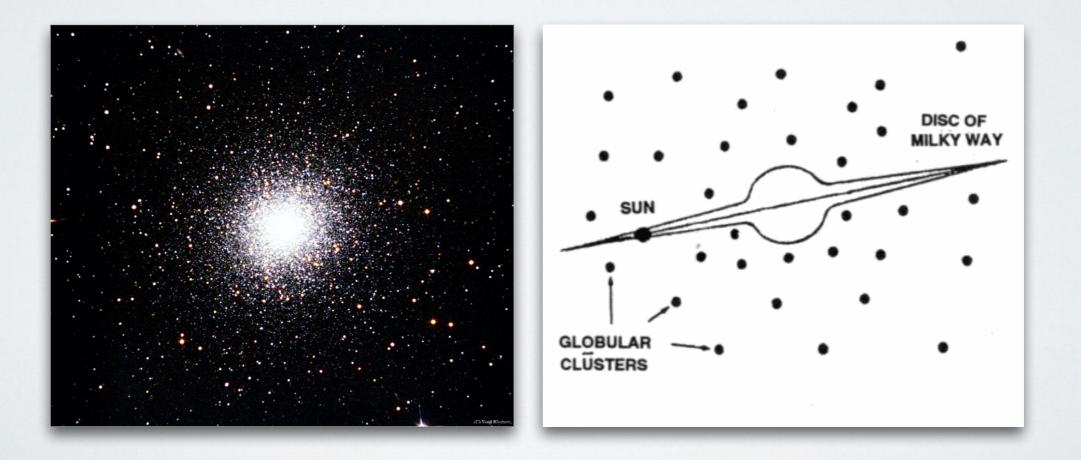


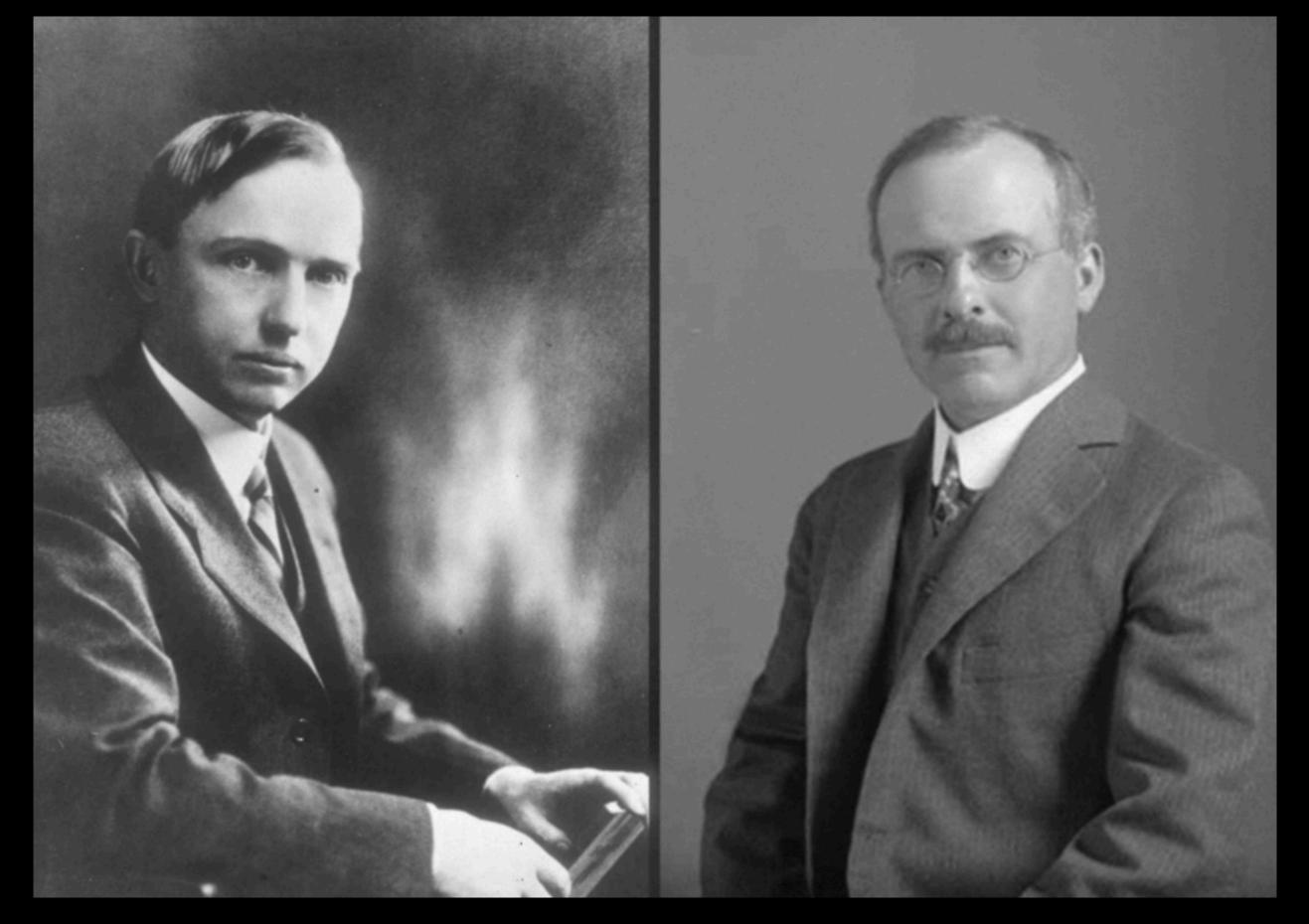
"Spiral nebula" Messier 51 by George Ritchey, 1910



Additional data about scale of the Universe

- Bessel measured **stellar parallax** of 61 Cygni (1838)
- Heber Curtis (1917) observed novae (strong flares in stars' light) in three spiral nebulae; much fainter than in MW, suggesting **great distance**
- Harlow Shapley showed that Sun is far out in disk of Milky Way (but overestimated MW size by factor of three)





Harlow Shapley

Heber Curtis

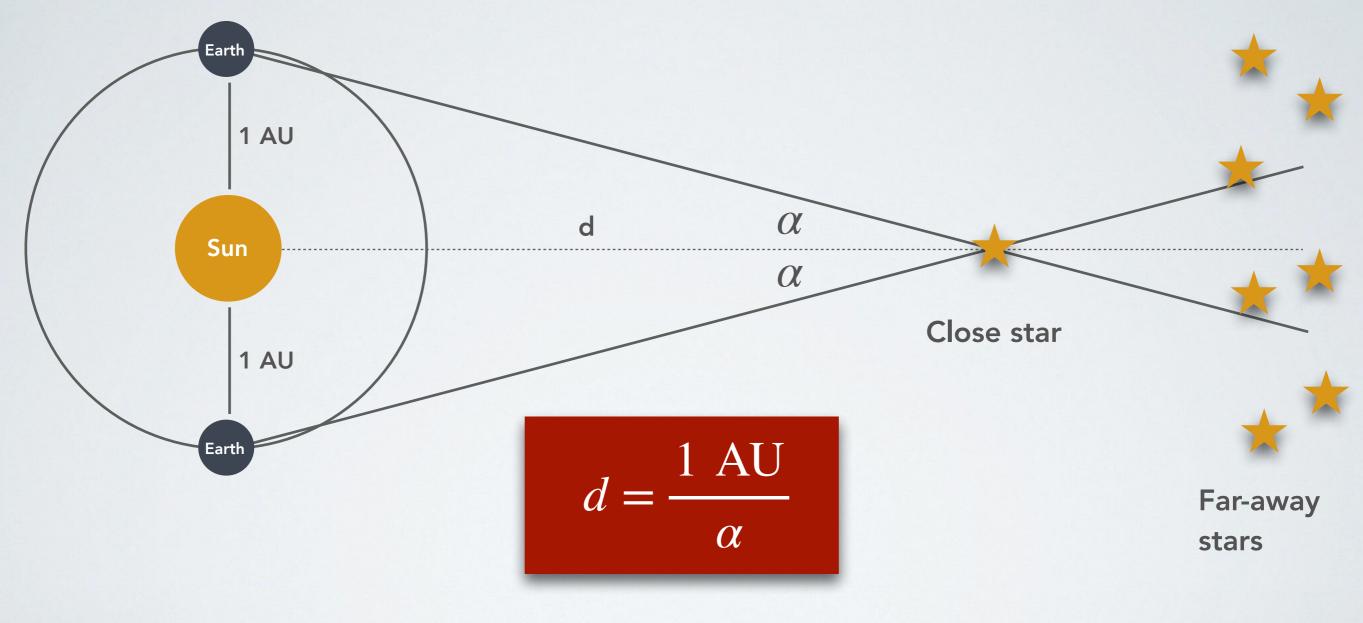
The Great Debate

- Shapley
 - If Andromeda not in the MW, must be extremely far away
 - Pinwheel galaxy seen to be rotating (wrong)
- Curtis
 - More novae in Andromeda than in MW
 - Novae appear faint in Andromeda (and other nebulae)



Part 3: The distance ladder

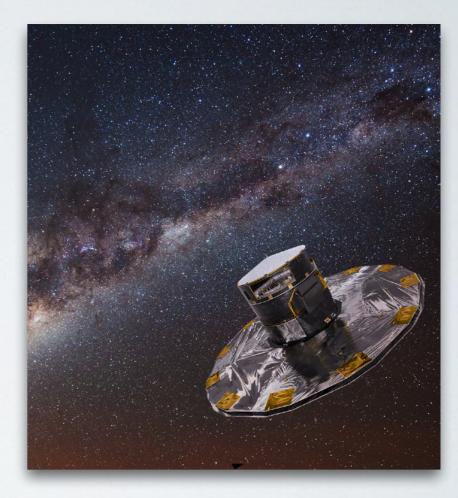
Parallax

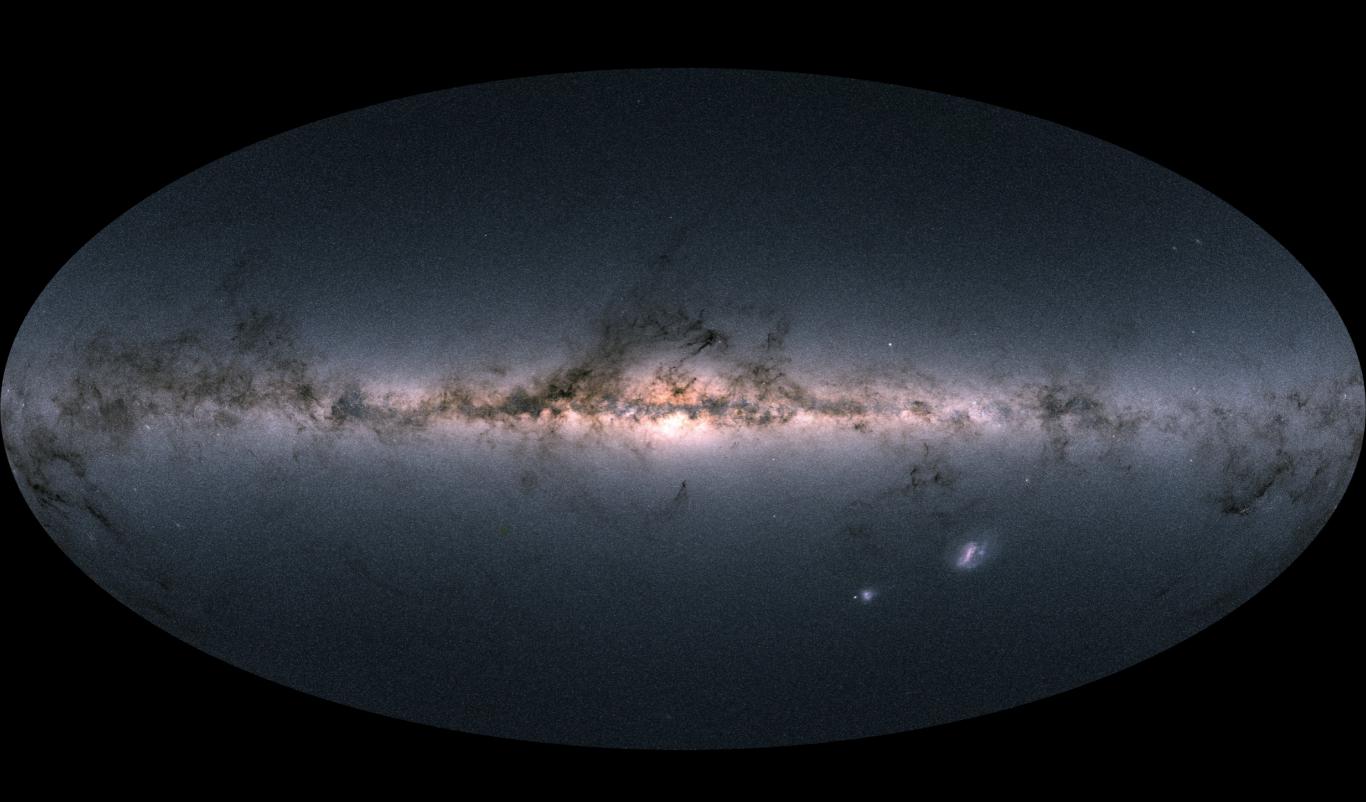


- Angles are expressed in radians
- sin(small angle) ~ small angle

Parallax

- Stellar parallax first measured in 1838 by Friedrich Bessel (0.29" for the star 61 Cygni)
- Until 1990s, could only detect parallax out to 50 pc
- GAIA satellite (launched 2013) can map out positions and motions of stars across the whole galaxy — 1 billion stars!







TurningPoint: How far away is the closest galaxy, Andromeda?



Parallax to other galaxies?

- Cannot detect parallax in galaxies outside the MW!
- Even for Andromeda (closest neighbor), parallax is

$$\theta = \frac{1''}{780,000} = 3 \times 10^{-10} \text{ degrees}$$

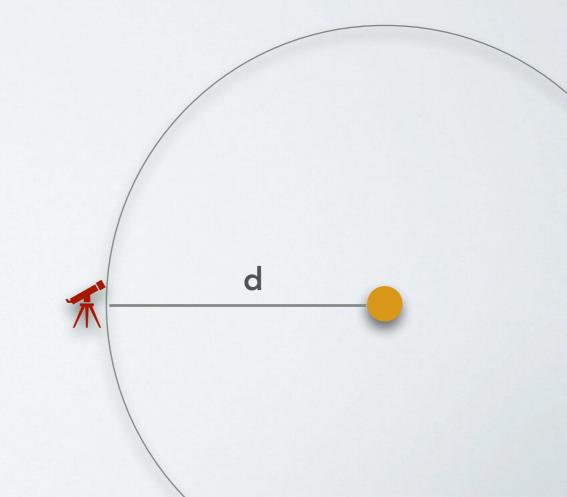


Standard Candles

- Fundamental issue: we cannot discern between an object being dim and being far away
- Need objects whose **absolute luminosity** we know; then:
 - Total luminosity is $L_{\rm std}$ (energy/time, e.g. erg/s or L_{\odot})
 - Observed brightness $b_{\rm obs}$ (energy/time/area, e.g. erg/s/cm²)
 - Distance is d, then

$$b_{\rm obs} = \frac{L_{\rm std}}{4\pi d^2}$$

$$\implies d = \sqrt{\frac{L_{\rm std}}{4\pi b_{\rm obs}}}$$



The first standard candle: Cepheid Variables



- In 1912, Henrietta Swan Leavitt observed a type of variable star called Cepheids
- Instrinsic luminosity can then be obtained from apparent brightness and **parallax distance**
- She discovered that Cepheids' total luminosity is related to the **period of fluctuations**
- Cepheids can be used as **standard candles!**

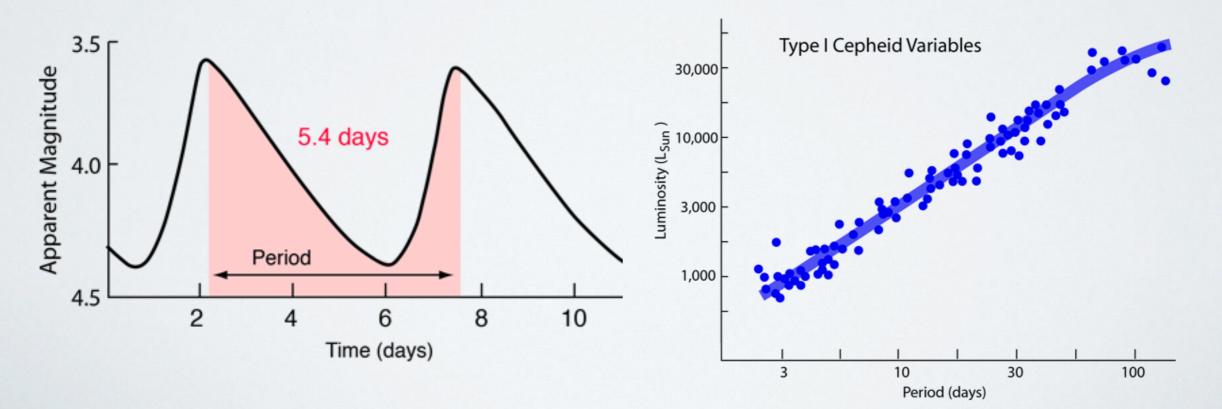
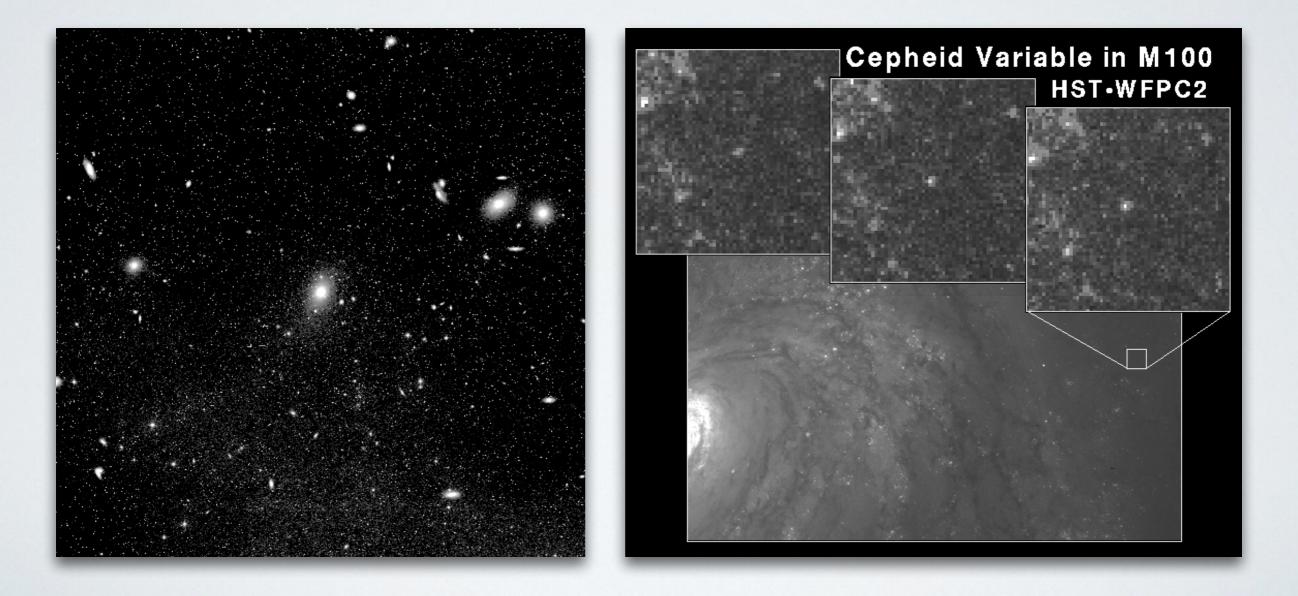


Image: hyperphysics.phy-astr.gsu.edu

Distance "ladder" normal star parallax distances Cepheid Cepheid distance

Cepheids Today

 In modern times, Cepheids in the Virgo galaxy cluster have been measured with Hubble Space Telescope (16 Mpc away!)



Participation: Group discussion #11

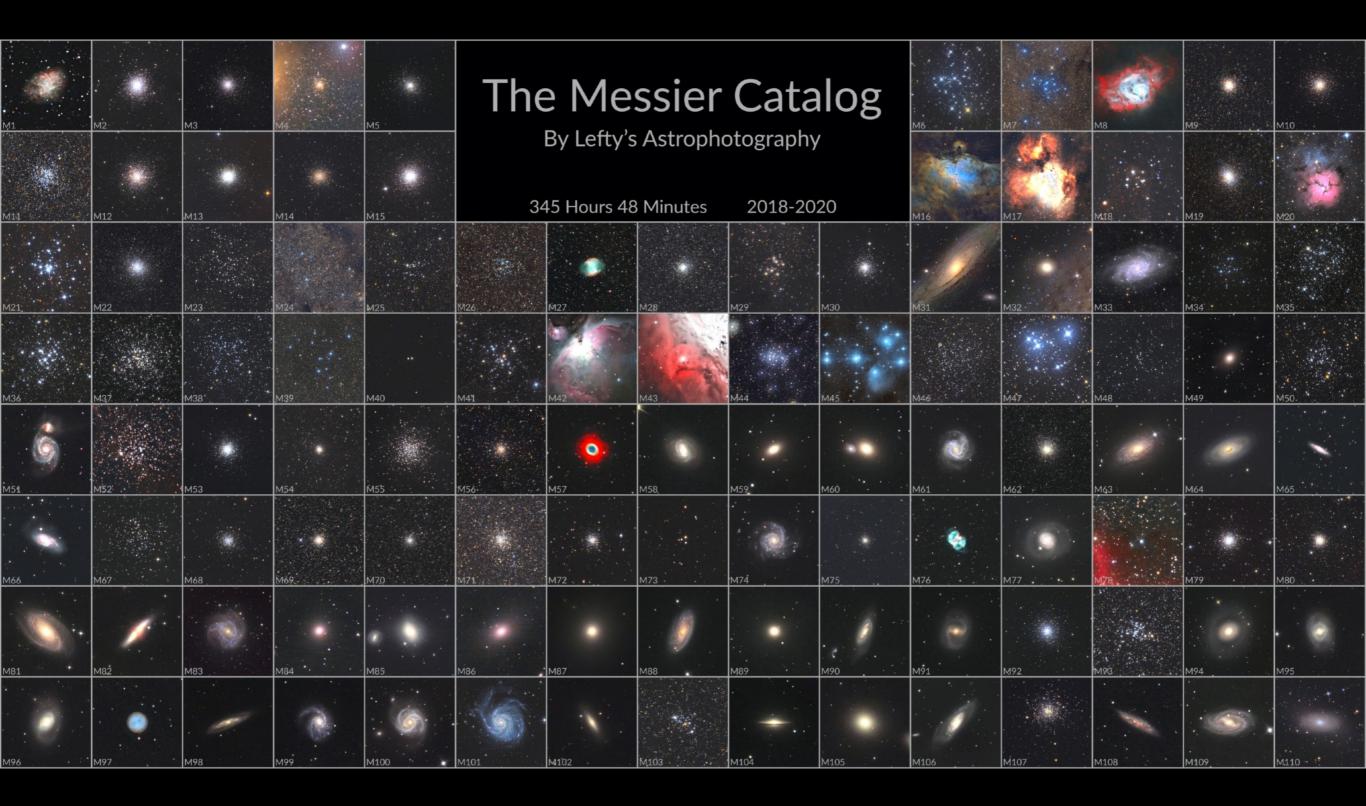


Messier Objects

Without knowing anything about what they are, how many categories would you sort them into?

Find a descriptive name for each category ("pink bunnies", "green squares"), write down a few of its characteristics, and list a few objects that would fall into this category. Are there any weirdos that don't fit any group?





Part 4: Hubble and the Expanding Universe

The distance to Andromeda



780 kpc



- Studied Andromeda "spiral nebula" with 100inch telescope on Mount Wilson
- In 1924, Edwin Hubble first observed a variable star with properties of a **Cepheid in Andromeda**

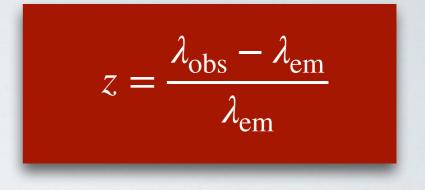
Andromeda

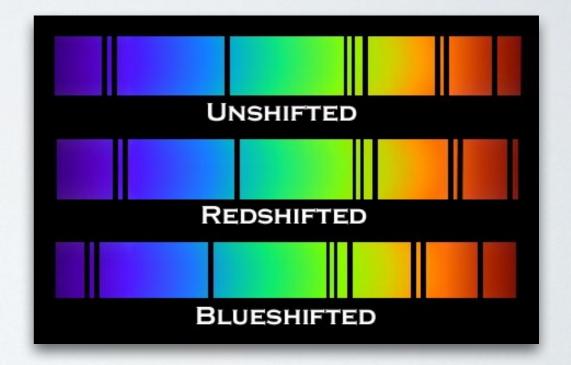
• Determined that Andromeda must be **well outside MW**, settling the Great Debate!



Cosmological Redshift

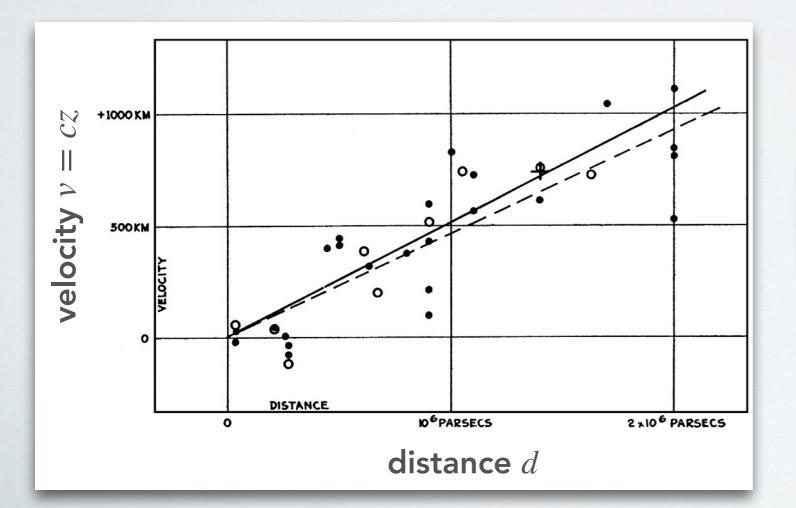
- Defined as the **relative shift in wavelength** between the emitted and observed light
- The "redshift" can be...
 - z > 0: redder, longer wavelength
 - z < 0: bluer, shorter wavelength
- Approximation z ≈ v/c can be used at low redshift and small velocities (v << c)





Hubble-Lemaitre law

- Slipher (1912) measured redshifts of some spiral nebulae, found large velocities (>1000 km/s) relative to MW
- Hubble and Humason systematically studied galaxies
 - Obtained redshifts from stellar spectra
 - Obtained **distances** using Cepheids and other estimates
- Interpreted redshift as Doppler shift, v/c (valid at low redshift)
- Linear relationship!
- Published in 1929; Lemaitre published same result in 1927 (in low-impact journal)



Vesto Slipher



Milton Humason

 $v = H_0 \times d$

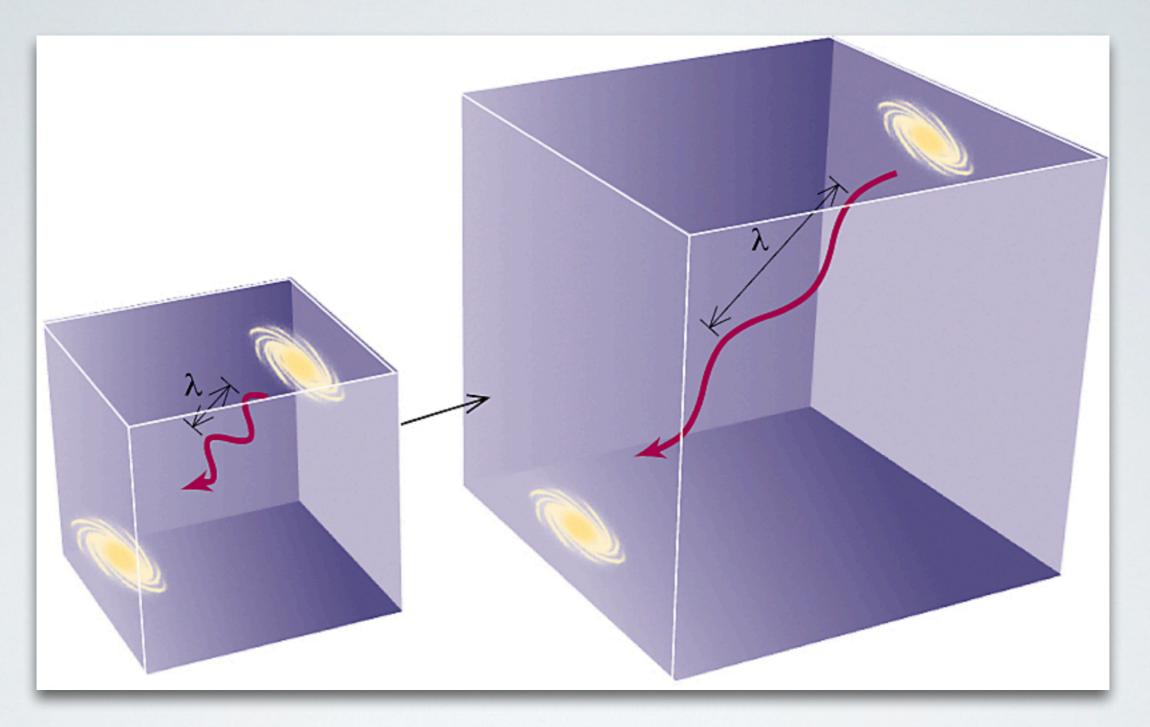
 $H_0 \approx 70$

km/s

Mpc

Georges Lemaitre

Cosmological redshift



Redshift is caused by the expansion of space!

Arny's explorations

Hubble & The basketball



Images: UChicago / NASA

Take-aways

- The enlightenment era was the beginning of extragalactic astronomy, but it was unknown whether galaxies ("nebulae") were inside Milky Way or separate objects
- We can use the distance ladder technique to calibrate standard candles that can be seen from farther away
- Galaxies that are **farther** away **move** away from us **faster**

Next time...

We'll talk about:

• Big Bang Theory & The expanding Universe

Assignments

- Post-lecture quiz (by tomorrow night)
- Homework #2 (by Thursday)

Reading:

• H&H Chapter 11